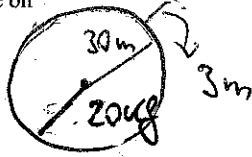


3E. One of the *Echo* satellites consisted of an inflated spherical aluminum balloon 30 m in diameter and of mass 20 kg. Suppose a meteor having a mass of 7.0 kg passes within 3.0 m of the surface of the satellite. What is the magnitude of the gravitational force on the meteor from the satellite at the closest approach? ssm

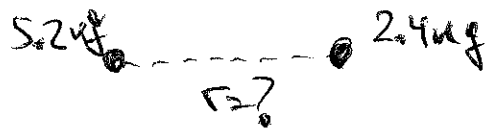


$F = ?$

3. We use $F = Gm_s m_m / r^2$, where m_s is the mass of the satellite, m_m is the mass of the meteor, and r is the distance between their centers. The distance between centers is $r = R + d = 15 \text{ m} + 3 \text{ m} = 18 \text{ m}$. Here R is the radius of the satellite and d is the distance from its surface to the center of the meteor. Thus,

$$F = \frac{(6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2 / \text{kg}^2)(20 \text{ kg})(7.0 \text{ kg})}{(18 \text{ m})^2} = 2.9 \times 10^{-11} \text{ N}$$

What is r if $F_g = 2.3 \times 10^{-12} \text{ N}$



1. The magnitude of the force of one particle on the other is given by $F = Gm_1m_2/r^2$, where m_1 and m_2 are the masses, r is their separation, and G is the universal gravitational constant. We solve for r :

$$r = \sqrt{\frac{Gm_1m_2}{F}} = \sqrt{\frac{(6.67 \times 10^{-11} \text{ N}\cdot\text{m}^2/\text{kg}^2)(5.2 \text{ kg})(2.4 \text{ kg})}{2.3 \times 10^{-12} \text{ N}}} = 19 \text{ m} .$$

$$F = \frac{G \cdot m_1 \cdot m_2}{r^2} \rightarrow r = \sqrt{\frac{G \cdot m_1 \cdot m_2}{F}}$$

An object is in equilibrium when positioned 150000 km from star X and 80000 km from star B. What is the ratio of masses (M_x/M_Y) of the two stars ?

$$F_x = G M_x M_{obj} / R_x^2 = F_y = G M_Y M_{obj} / R_Y^2$$

$$G M_x M_{obj} / (150000 \cdot 10^3 \text{ m})^2 = G M_Y M_{obj} / (80000 \cdot 10^3 \text{ m})^2$$

$$(M_x/M_Y) = (80000 \cdot 10^3 \text{ m} / 150000 \cdot 10^3 \text{ m})^2$$

$$(M_x/M_Y) = 0.284$$

18. ORGANIZE AND PLAN Use Newton's law of gravitation (Eq. 9.1) to calculate the attractive force due to gravity between the two balls.

Known: $m_1 = 25 \text{ kg}$, $m_2 = 60 \text{ kg}$, $r = 0.50 \text{ m}$.

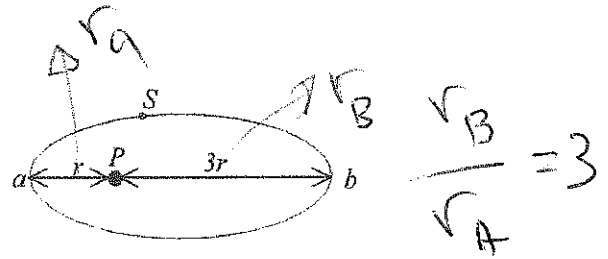
SOLVE Newton's law of gravitation gives [Eq. 1]

$$F = \frac{Gm_1m_2}{r^2} = \frac{(6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2/\text{kg}^2)(25 \text{ kg})(60 \text{ kg})}{(0.50 \text{ m})^2} = 4.00 \times 10^{-7} \text{ N}$$

This is response (d).

REFLECT The problem does not state the radii of the two balls involved. Does this matter?

A satellite moves around a planet in an elliptical orbit. What is the ratio of the speed of the satellite at point a to that at point b ?



Answer:

By the conservation of angular momentum, we have

$$I_a \omega_a = I_b \omega_b = L = \text{const} ; I = mr^2$$

where I_a and I_b are the moment of inertia of the system about P at positions a and b respectively.

$$v = \omega \cdot r \rightarrow \omega = \frac{v}{r}$$

Now, we can write $mr_a^2 \left(\frac{v_a}{r_a}\right) = mr_b^2 \left(\frac{v_b}{r_b}\right) \rightarrow r_a v_a = r_b v_b$

which gives $\frac{v_a}{v_b} = \frac{r_b}{r_a} = \frac{3}{1}$. That is, $v_a : v_b = 3 : 1$.